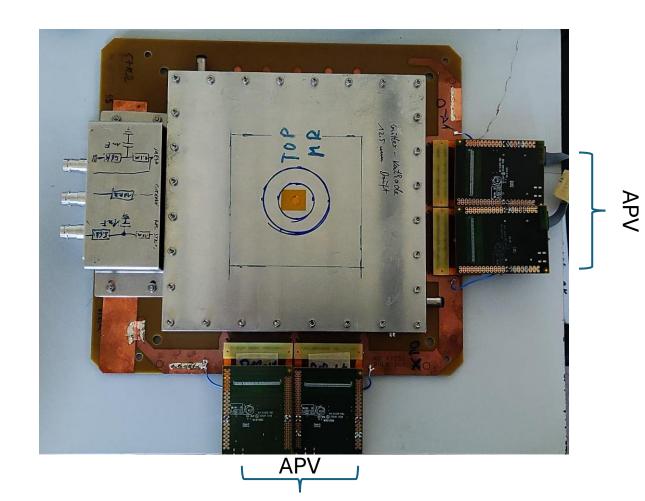
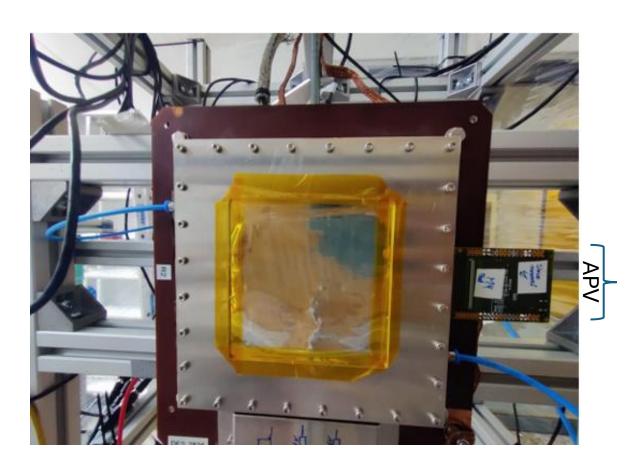
# Hardware Meeting 19.09.2024

## Motivation:

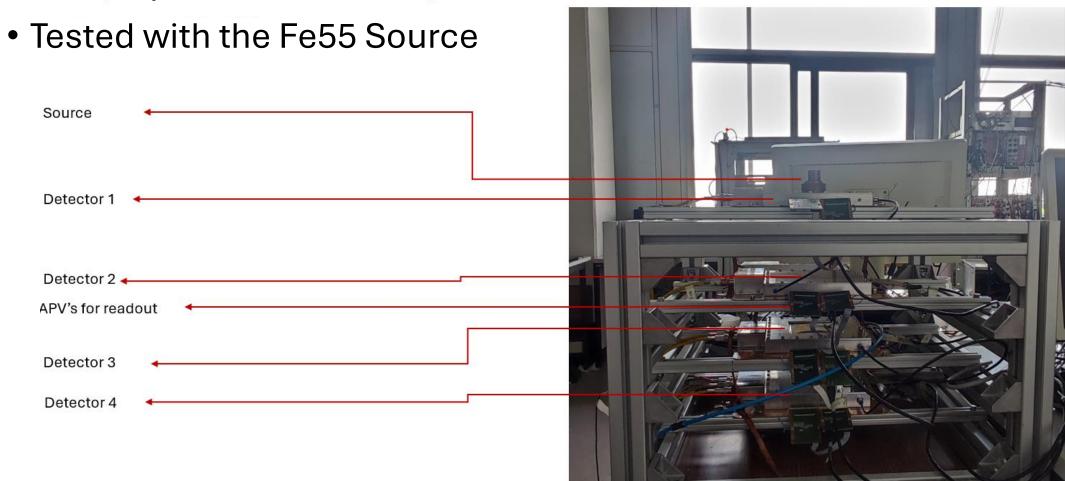
- Investigation of PAD Detectors.
  - Reduced read out channels and consequently lower number of electronics compared to conventional MicroMeGas.

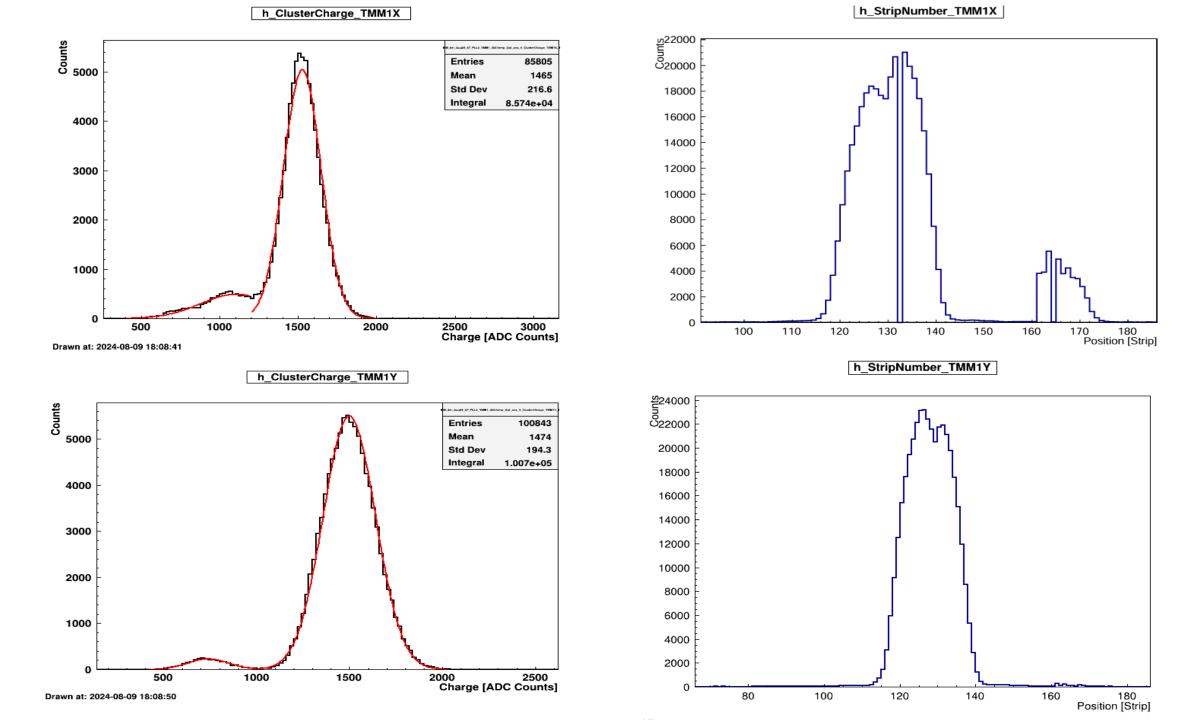




## Progress

 Tested the 4 TMM's to select which ones to use for the new telescope

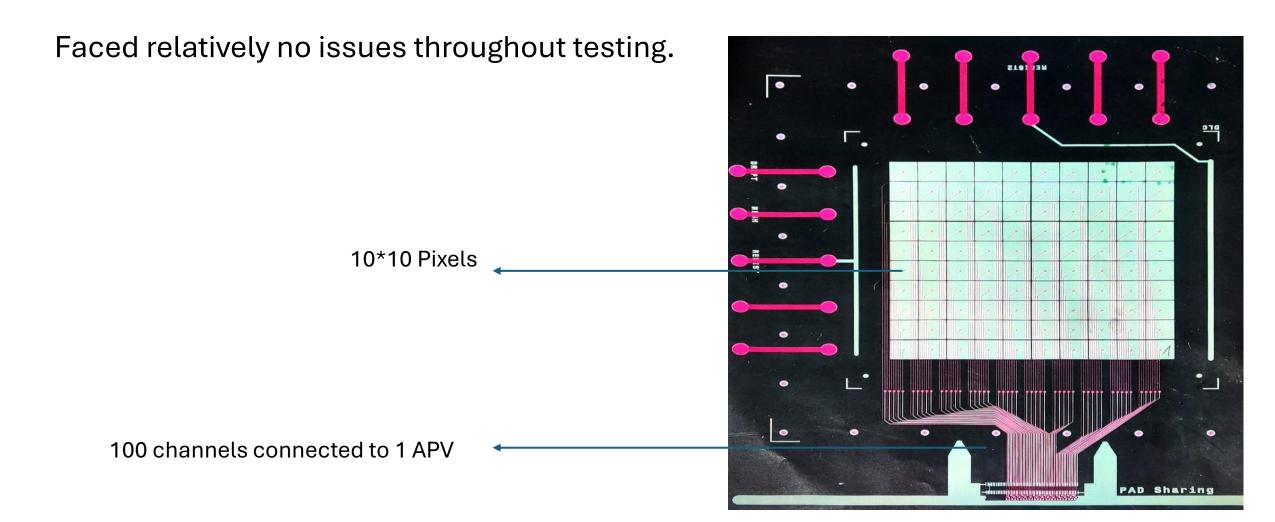


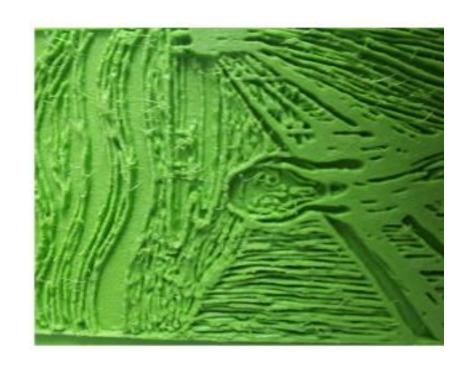


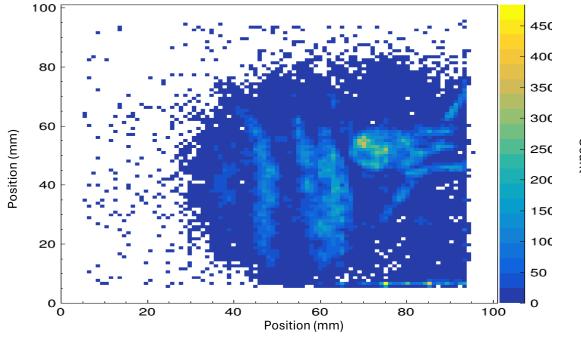
• PAD 2 Test:

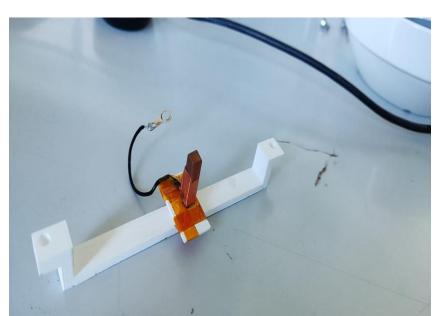
PAD2: Old pad detector with 5 layers

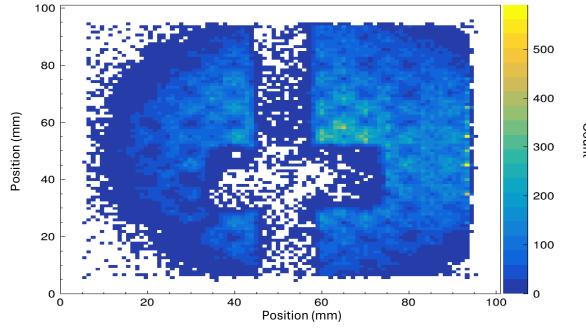
- 10\*10 Pixels, 1 APV and 100 readout channels.
- Tested using Fe55 source.
- Position Reconstruction tested with 3D printed Scream portrait and Cu-block.
- Pitch = 10mm







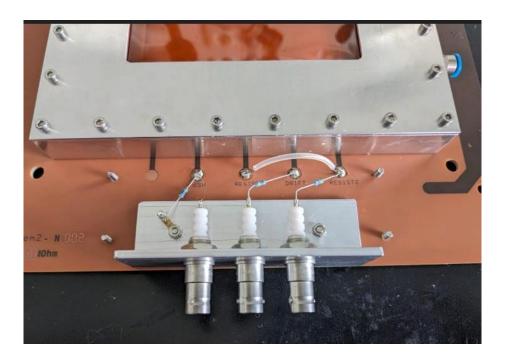




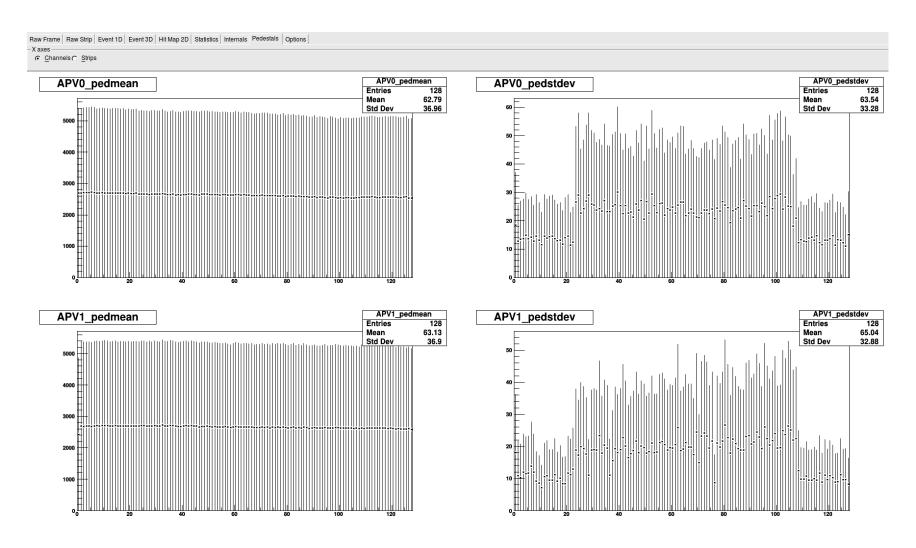
## • PAD 1 Test:

- PAD1: New pad detector with 3 layers
- 2 APV's

- New Unknown Pad Det.
- Issues Faced and current solutions:
  - Noise and grounding issues Oscilloscope signals were too noisy, and we could not identify the peaks.
    - 1. Initially the entire edges of the detector was covered using Cu tape.
    - 2. The detector was placed on an Aluminum plate. Helped a bit.
    - 3. Additional RC filter was installed before the timing filter amplifier. This reduced the noise by a huge amount.



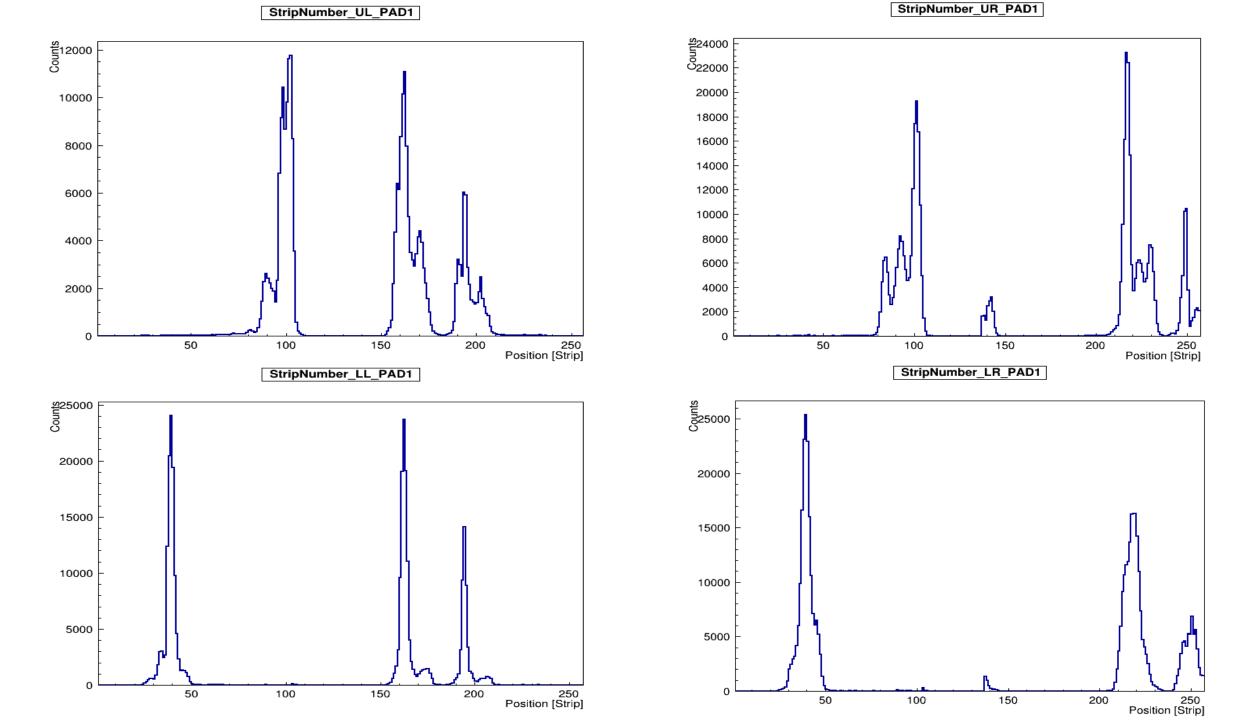
- Strip-detector like behavior.
  - 1. Looked at mmdaq on etp10 to see the connected channels.



• Different rates were observed across the detector surface at the same voltage under the same conditions with the Fe55 source.

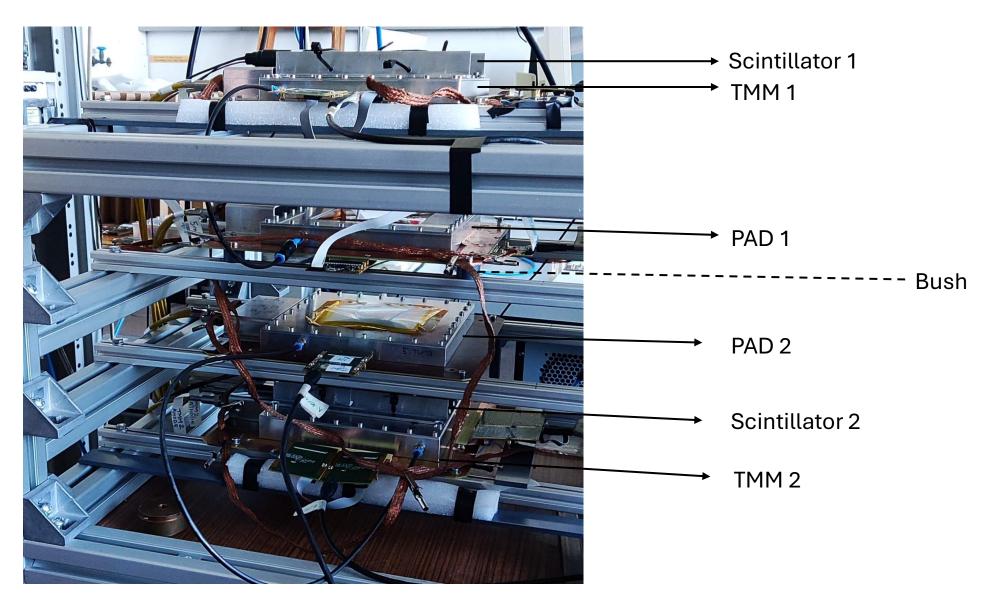
To Voltage Supply		65Hz	570Hz	200Hz
		1000Hz	1100Hz	900Hz
		120Hz	600Hz	400Hz

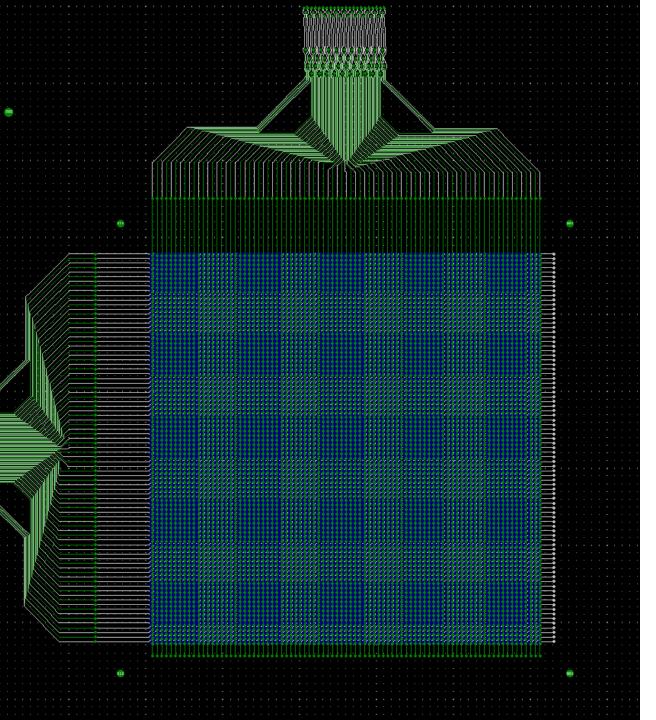
• Still don't have a solution to this problem.



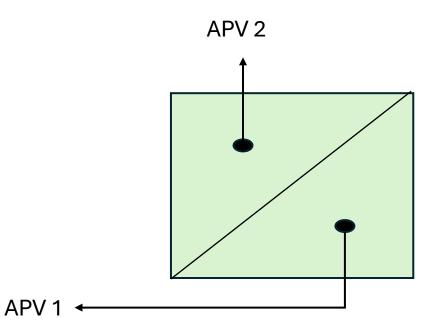
APV Connections below the detector which lead to problems with building the telescope setup.

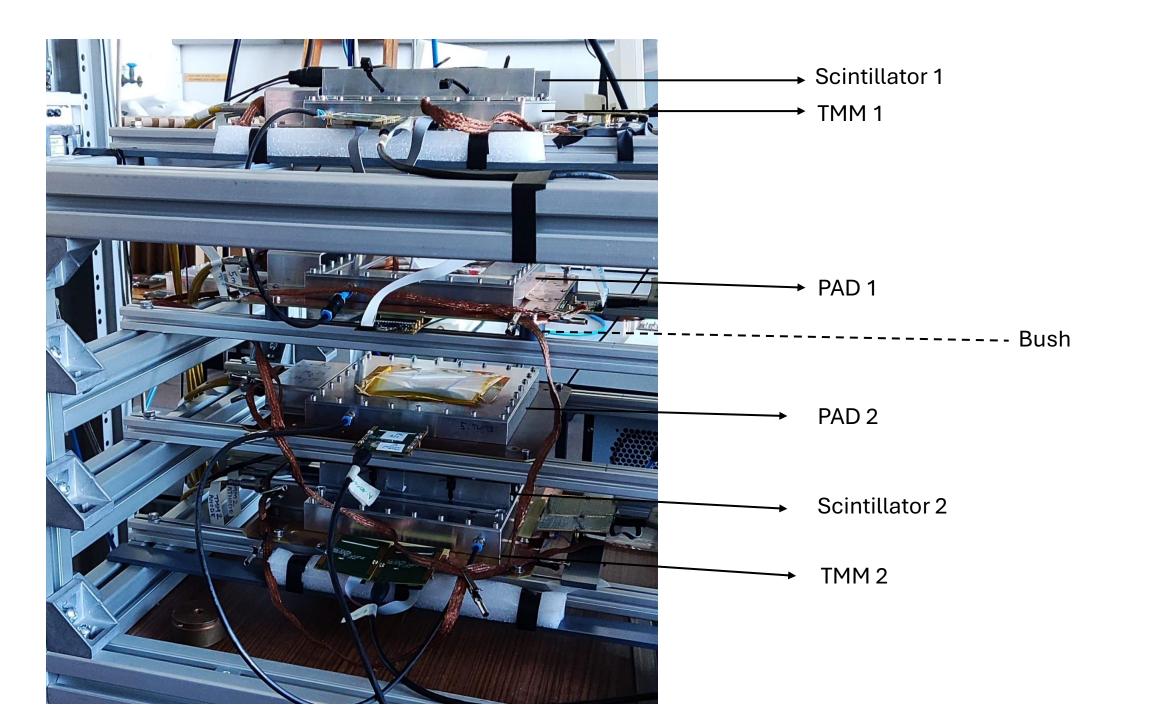
1. Rectified with a plastic bush.

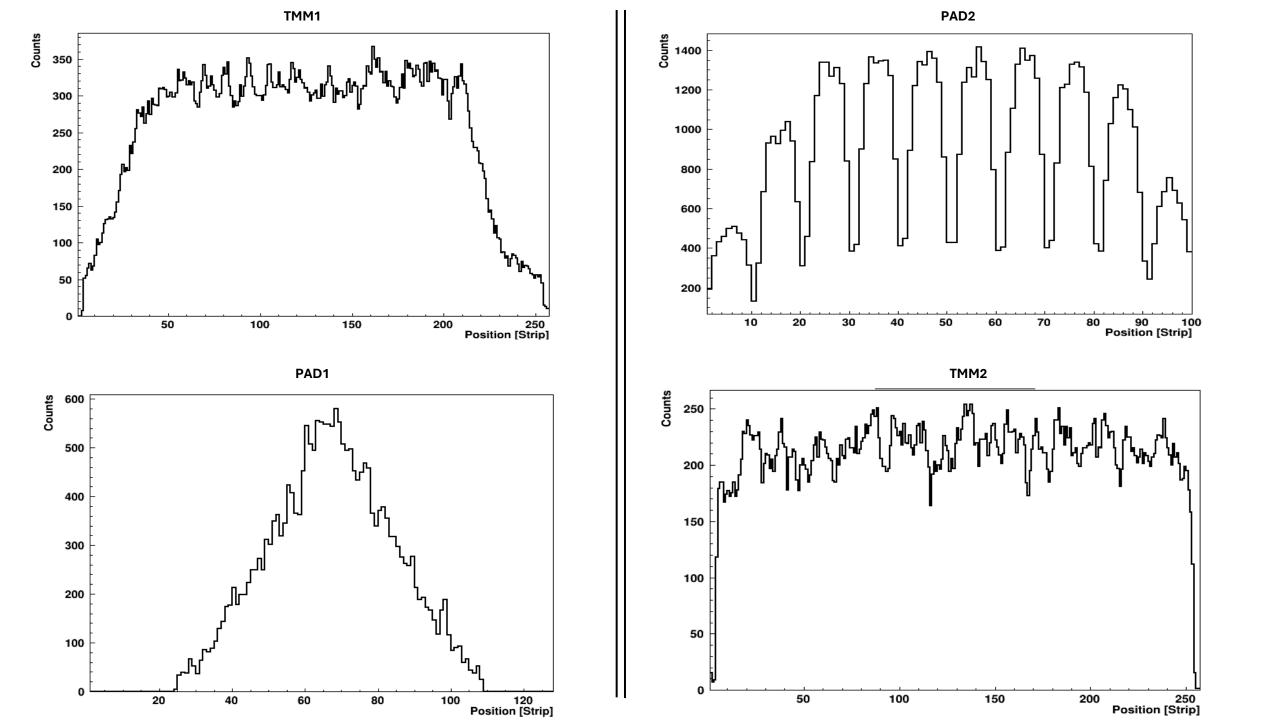




- Investigation of PAD 1:
  - The strip like behavior of PAD led us to pixel shape and connections to APV.
  - 3 layers
  - 85\*85 pixels (3<sup>rd</sup> layer)
  - 2 APV's
  - Pitch 1.2mm (Outermost pixel layer)
  - 170 channels



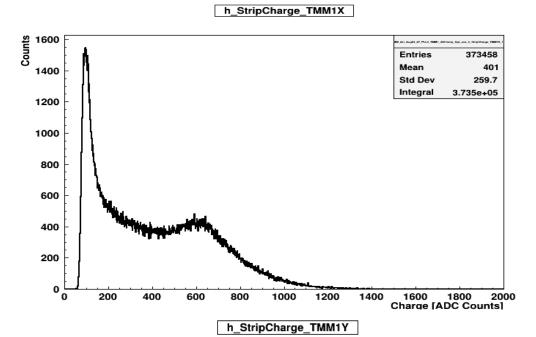


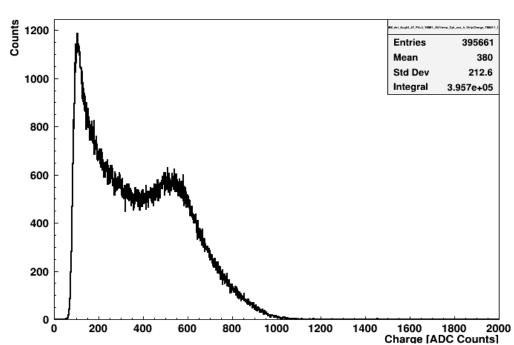


## What next?

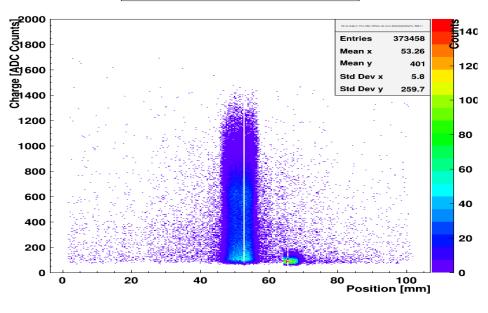
- Charge reconstruction for PAD1.
- Alignment of the detectors in the script.
- Tracking to find the Residuals for the telescope.
- Work to be done with Clustering for the pixel detectors.

### • TMM 1 additional Plots

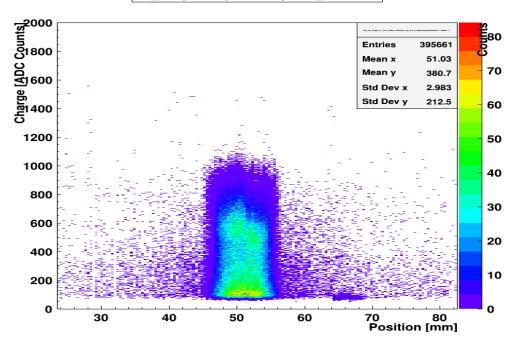




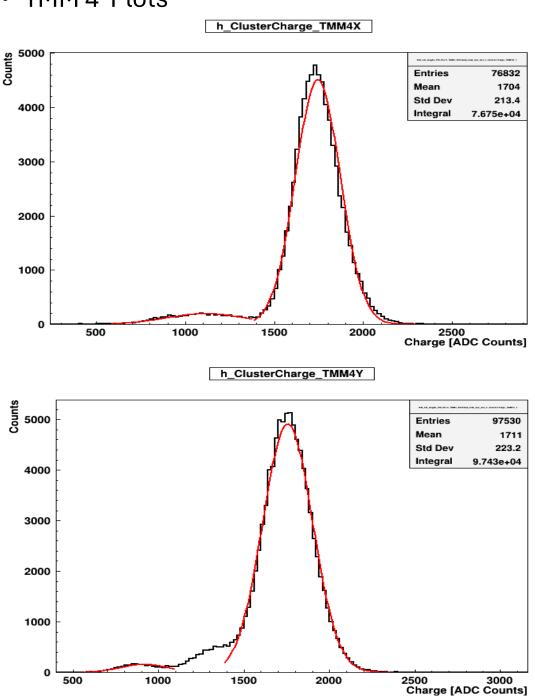
#### h\_StripChargeVSStripPos\_TMM1X

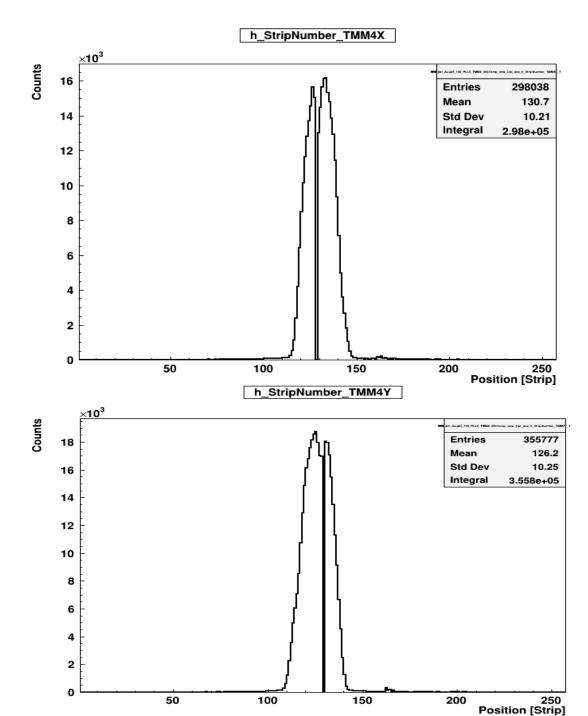


#### h\_StripChargeVSStripPos\_TMM1Y

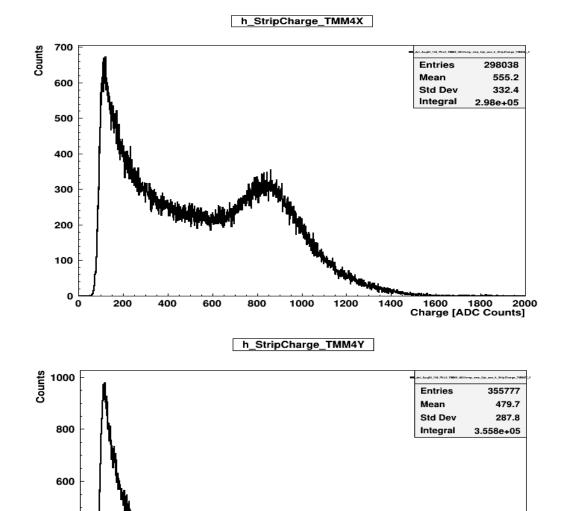


### • TMM 4 Plots



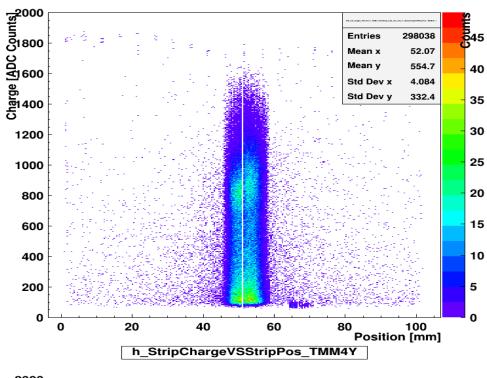


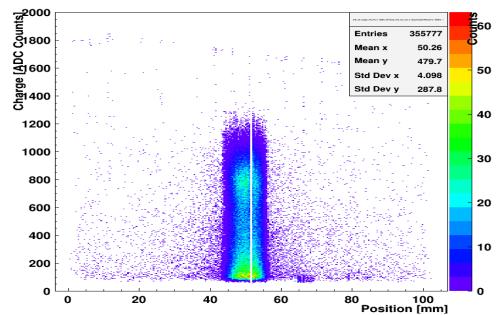
## • TMM 4 additional Plots



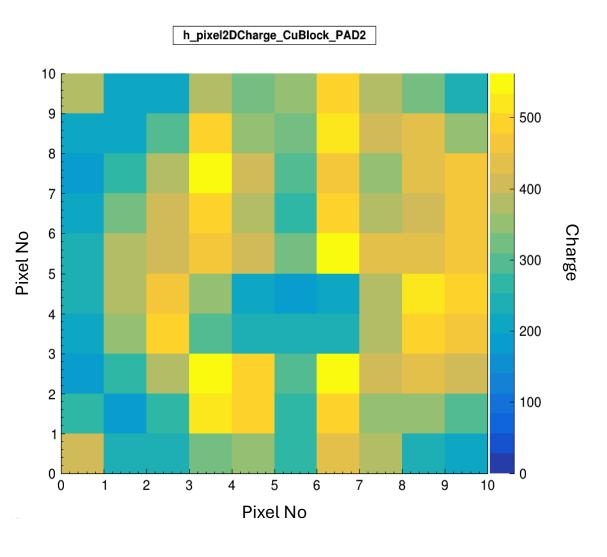
1600 1800 200 Charge [ADC Counts]

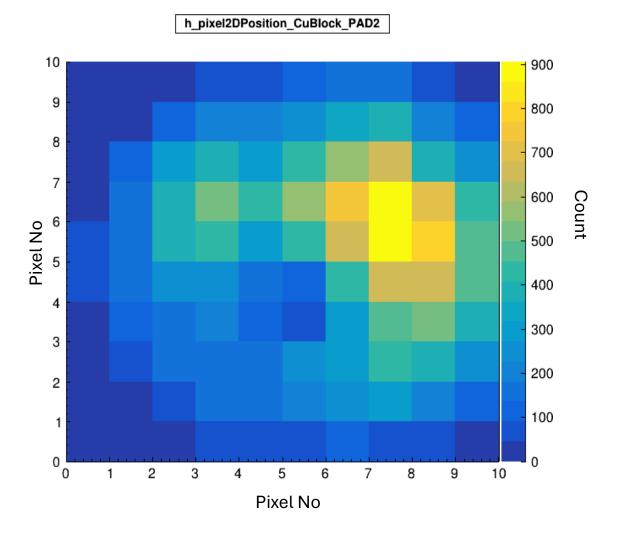
#### h\_StripChargeVSStripPos\_TMM4X

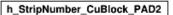


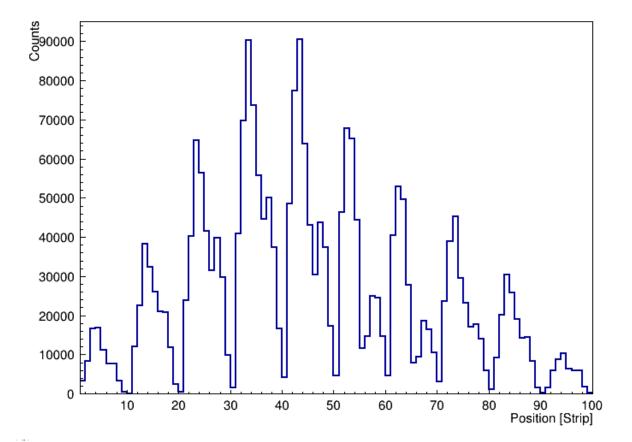


## PAD2 additional Plots

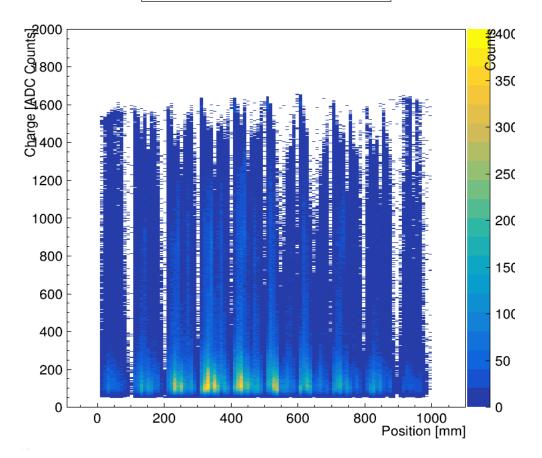


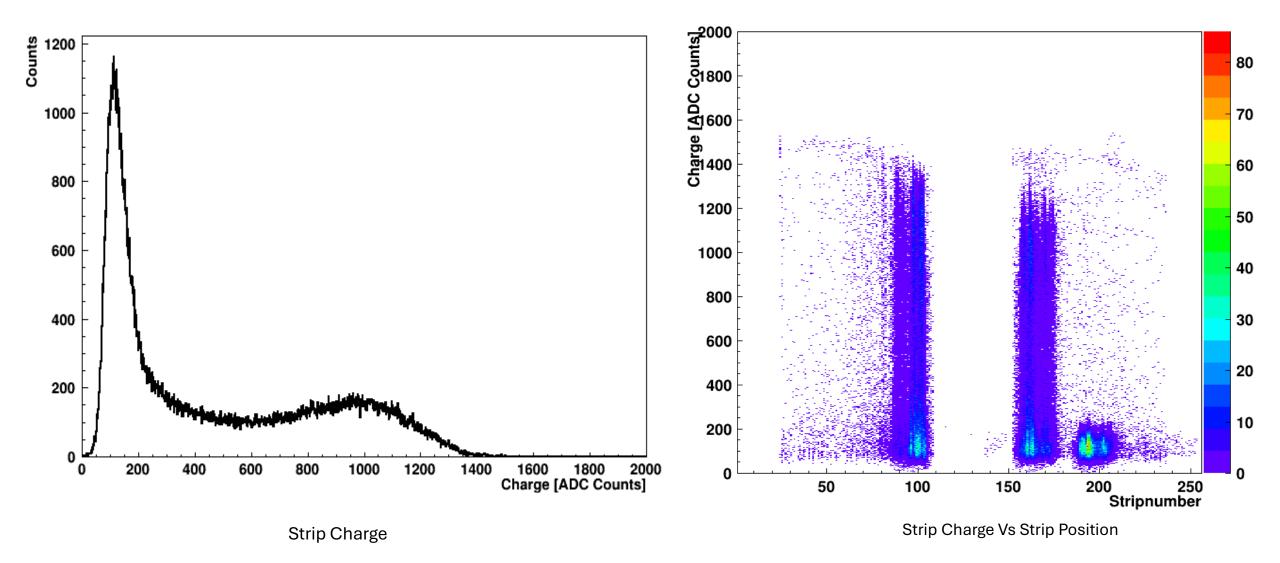


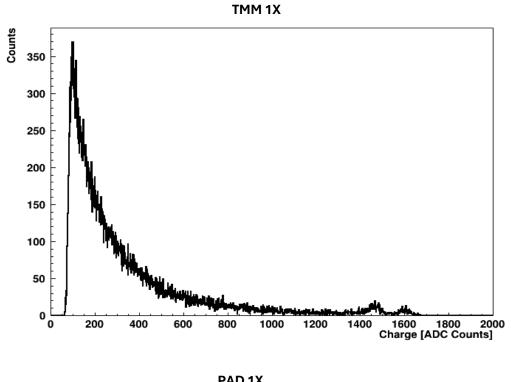


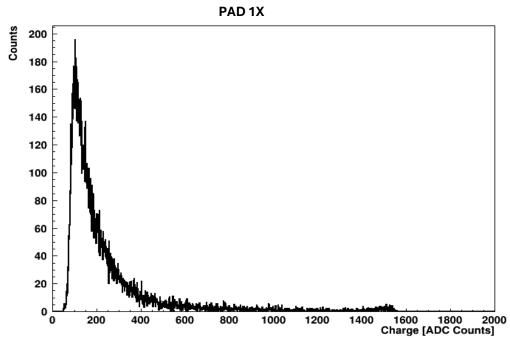


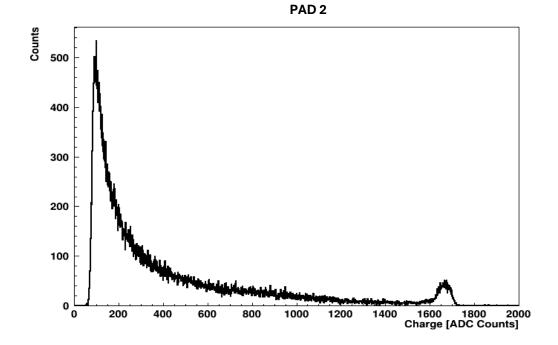
#### h\_StripChargeVSStripPos\_CuBlock\_PAD2

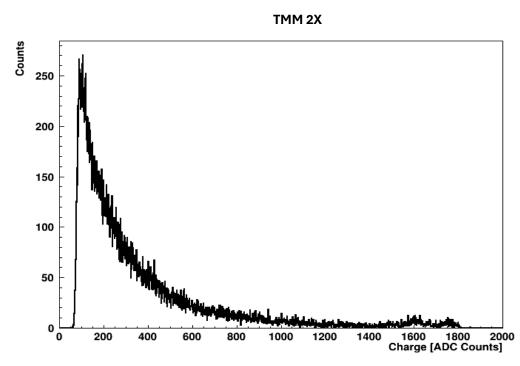


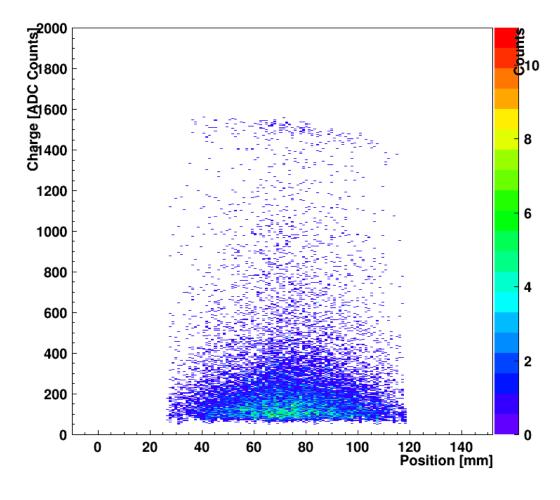












Counts 16 18 20 Charge [ADC Counts] 

Strip Charge Vs Strip Position PAD 1X

Cluster Charge PAD 1X