

Area C Overview

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IDT-UM Meeting
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CLUSTER OF EXCELLENCE
QUANTUM UNIVERSE

Introduction

Reminder: Structure of Area C

Deep Learning, Erkenntnisgewinn durch fundierte datengetriebene Methoden

<p>C1) Sensornahe Verarbeitung von Daten</p> <ul style="list-style-type: none">• Signalfilter, Rauschunterdrückung• Verarbeitung von zeitabhängigen Signalen	<p>C2) Objektrekonstruktion</p> <ul style="list-style-type: none">• Spur- und Clusterrekonstruktion, Jetbildung, Ereignisrekonstruktion• Fragestellungen für Anordnung, Reihenfolge, Zuordnungen von Daten• Optimierungen zur Extraktion kleiner Signale bei großem Untergrund
<p>C3) Netzwerkbeschleunigte Simulationen</p> <ul style="list-style-type: none">• Generative adversarial networks, Anpassung von Simulationen an Datenverteilungen• Evaluationsverfahren für die Qualität der Netzwerksimulationen	<p>C4) Qualität von Netzwerkvorhersagen</p> <ul style="list-style-type: none">• Reduzierung experimenteller systematischer Unsicherheiten• Spezielle Lernstrategien• Vorhersagenrelevante Information• Unsicherheiten von Vorhersagen

Updates

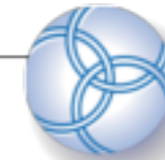


C1) Sensornahe Verarbeitung von Daten

- Signalfilter, Rauschunterdrückung
- Verarbeitung von zeitabhängigen Signalen

C2) Objektrekonstruktion

- Spur- und Clusterrekonstruktion, Jetbildung, Ereignisrekonstruktion
- Fragestellungen für Anordnung, Reihenfolge, Zuordnungen von Daten
- Optimierungen zur Extraktion kleiner Signale bei großem Untergrund



FIAS Frankfurt Institute for Advanced Studies



ERLANGEN CENTRE FOR ASTROPARTICLE PHYSICS



C3) Netzwerkbeschleunigte Simulationen

- Generative adversarial networks, Anpassung von Simulationen an Datenverteilungen
- Evaluationsverfahren für die Qualität der Netzwerksimulationen

C4) Qualität von Netzwerkvorhersagen

- Reduzierung experimenteller systematischer Unsicherheiten
- Spezielle Lernstrategien
- Vorhersagenrelevante Information
- Unsicherheiten von Vorhersagen



Universität Hamburg
DER FORSCHUNG | DER LEHRE | DER BILDUNG



Karlsruher Institut für Technologie

**Fast processing of data
close to sensors (CI)**

Reminder: Aim of the project in Mainz

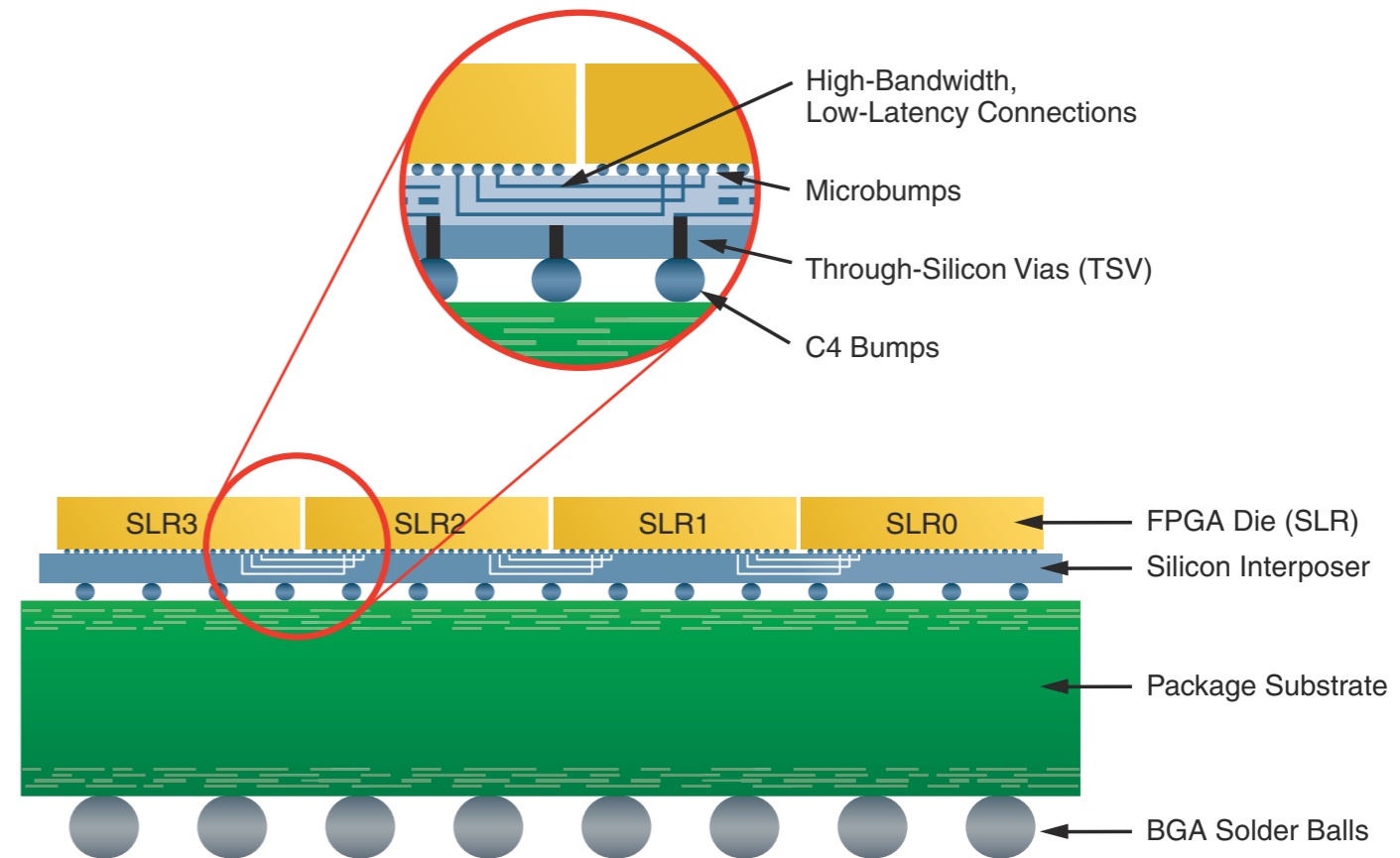
- **Processing of detector data at extremely high rates**
 - Not possible to store data due to its size
 - Usage of GPUs not possible due to their too high latency
 - Data has to be processed and filtered locally in real-time, maybe directly at the corresponding sensors
- **Solution: deep neural networks** as replacement for iterative algorithms, that can be **efficiently evaluated on FPGAs**
- **Test environment: ATLAS L1 Trigger** (40 MHz event rate)

Current status

- **First implementation of Dense, 2D-Convolution and MaxPooling layer done**
 - Paper on implementation details as well as performance evaluation published: [JINST 14, P09014 \(2019\)](#)
(implementation only, assumes data already on FPGA)
- Ongoing firmware development to cope with technical details of modern large FPGAs
 - **Not monolithic chips** but several individual chips (i.e. 3 for our reference FPGA from Xilinx), *see next slide for details*
 - **Data IO** to and from FPGA via optical fibres and multi-gigabit transceivers (MGTs) **non-trivial** for our test case (different link speeds need to be handled simultaneously)

Modern FPGA: the devil is in the details ...

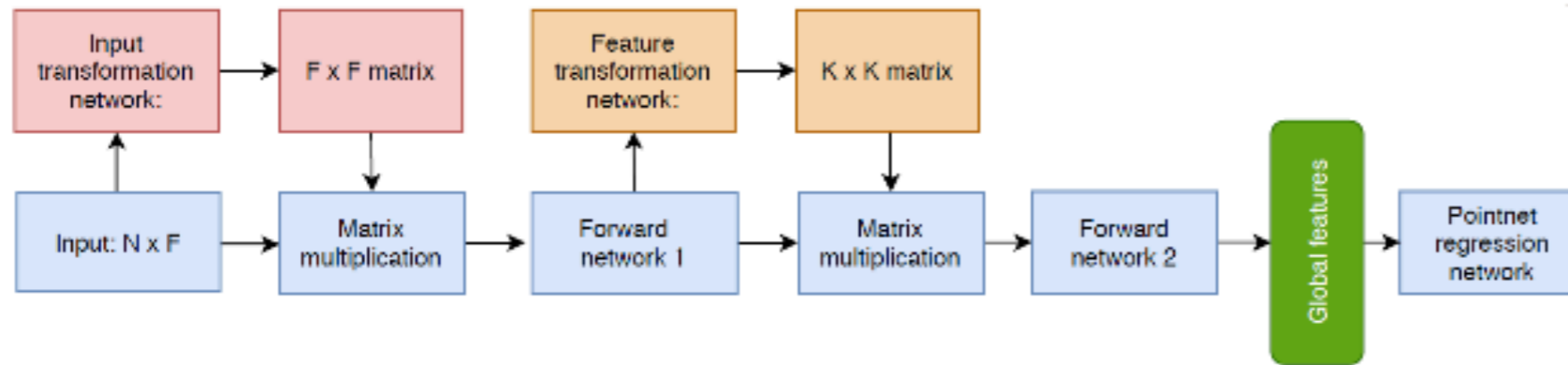
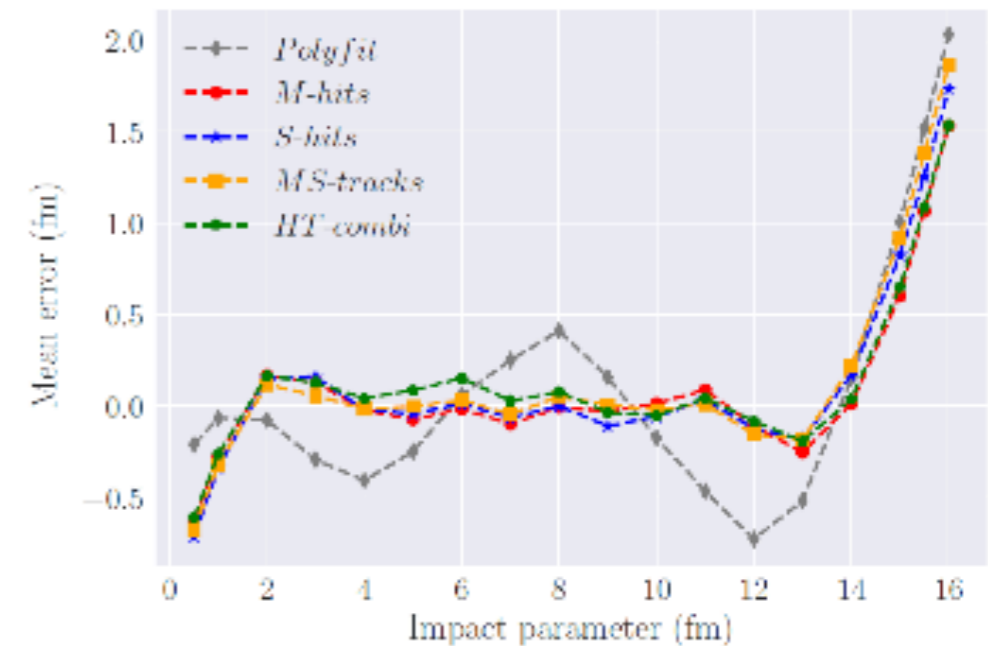
- **Several chips connected with finite connection**
 - Potential bottleneck depending on inputs and network architecture (only ~17k inter-chip connections)
- **Data input distributed over all SLRs, especially problematic for larger convolution layers at the start of the network**
 - **Routing** via design tool (Xilinx Vivado) **becomes challenging** once resource usage increases (larger networks)



Object reconstruction (C2)

DL analysis tools for CBM (FIAS team)

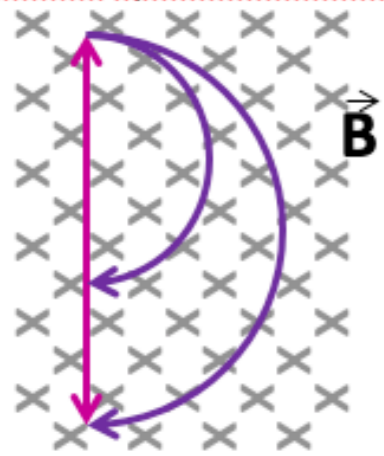
- Analysis of simulated CBM data using PointNet turned out to be accurate and fast
 - ~O(1000) events per GPU (NVIDIA) without optimizing for speed.
- Can do centrality selection or physics identification
 - M. O. Kuttan, J. Steinheimer, K. Zhou, A. Redelbach and H. Stoecker, [arXiv:2009.01584 [hep-ph]].
 - EoS classification : In preparation



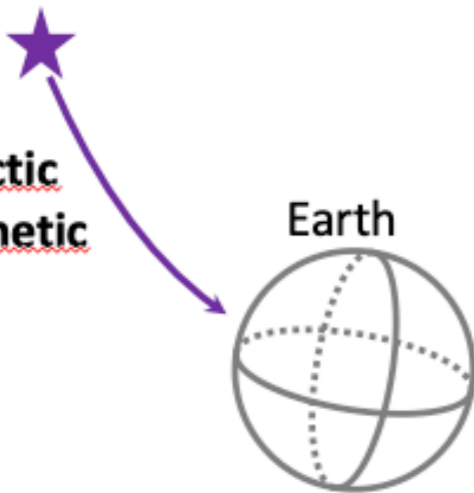
Prediction of impact parameters from unordered sets of tracks/hits

Patterns in cosmic ray arrival directions

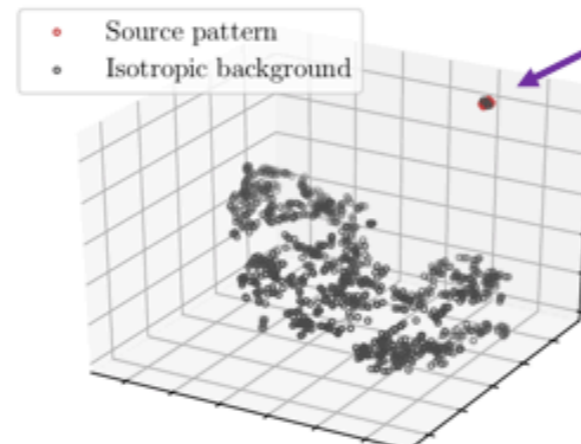
Mass spectrometer



Galactic
Magnetic
Field

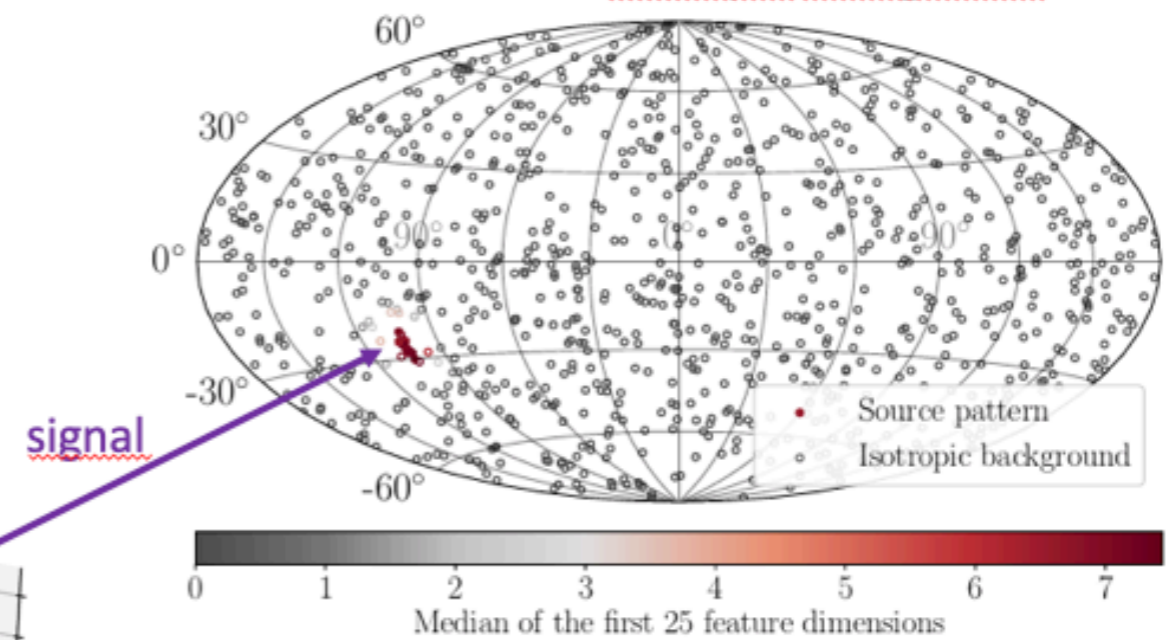


Graph network recognizes
cosmic ray arrival pattern,
separates it from
background cosmic rays

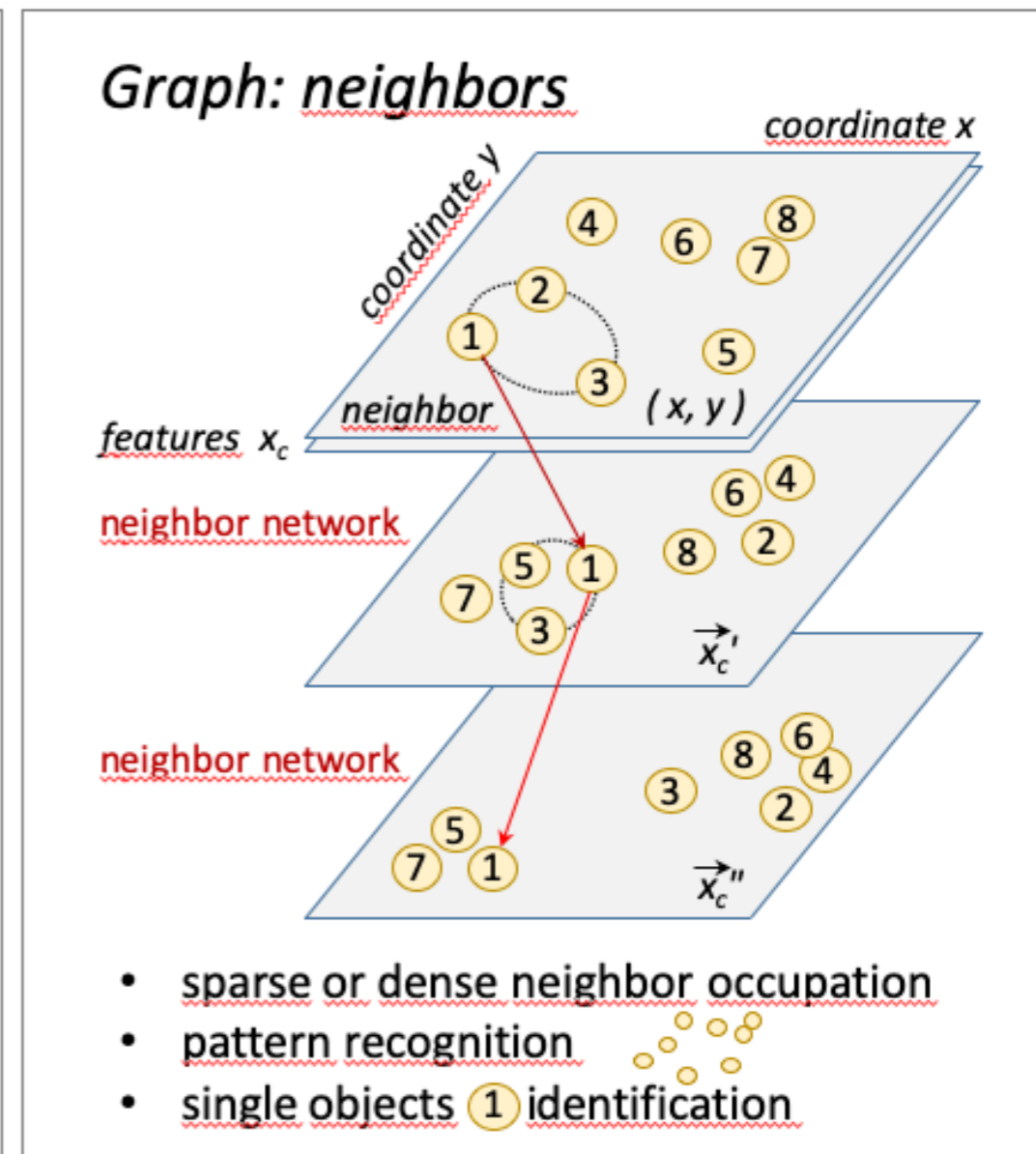
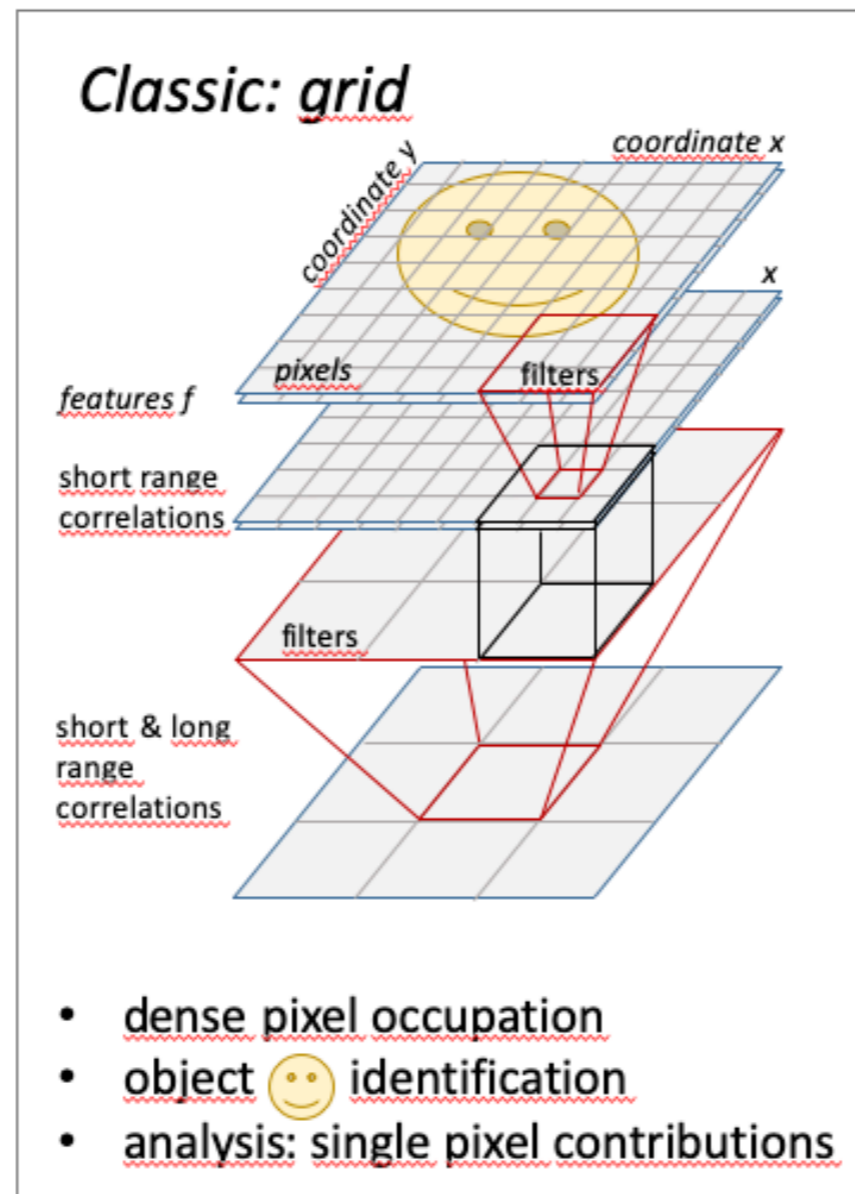
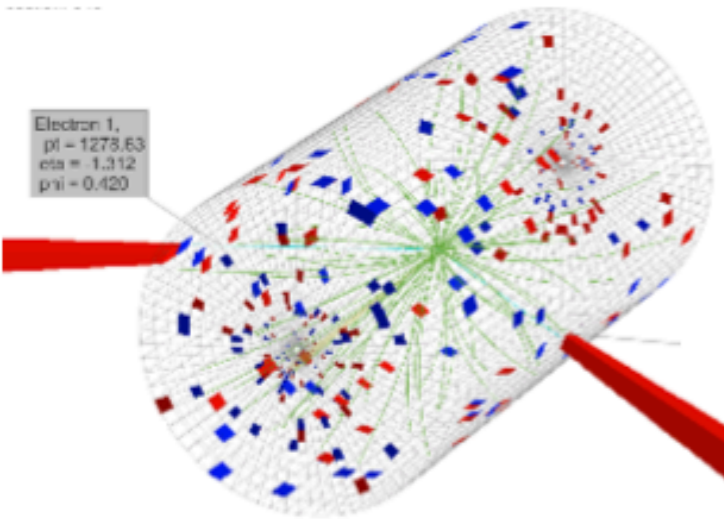


A three-dimensional representation of the final EdgeConv layer's output is created using the t-SNE technique [68].

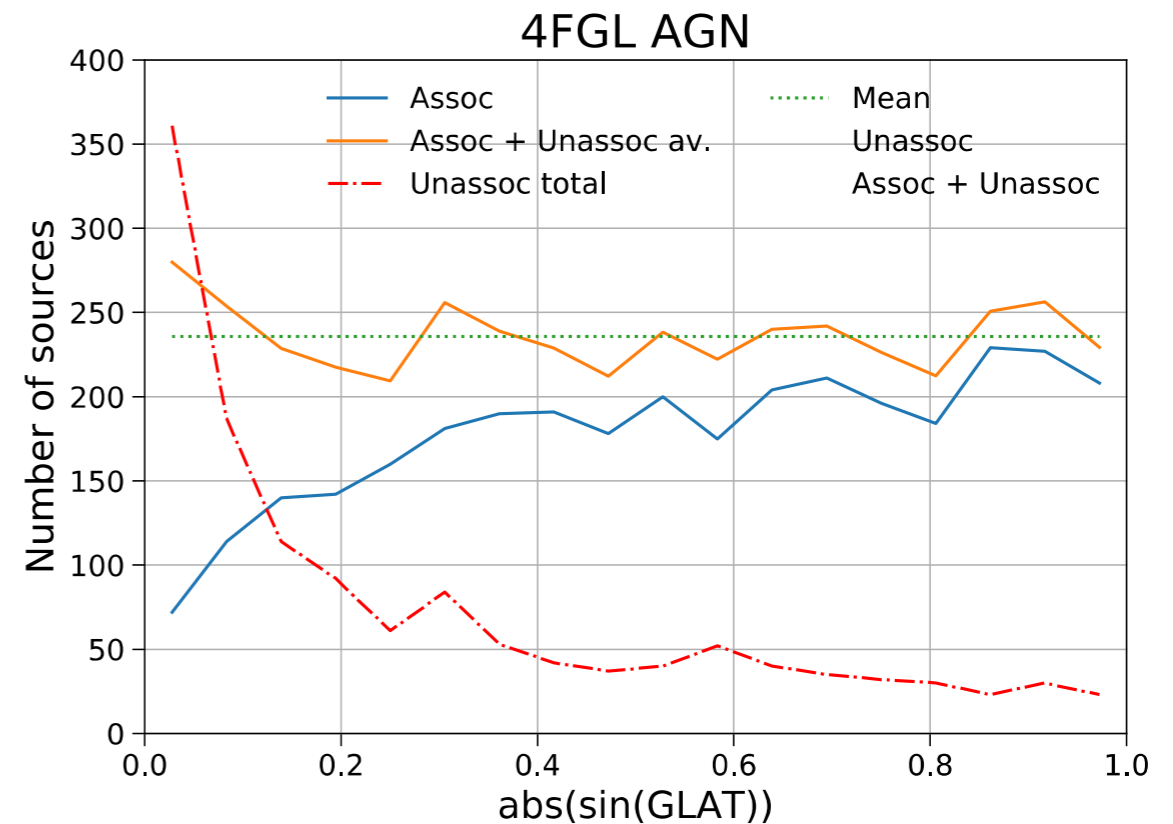
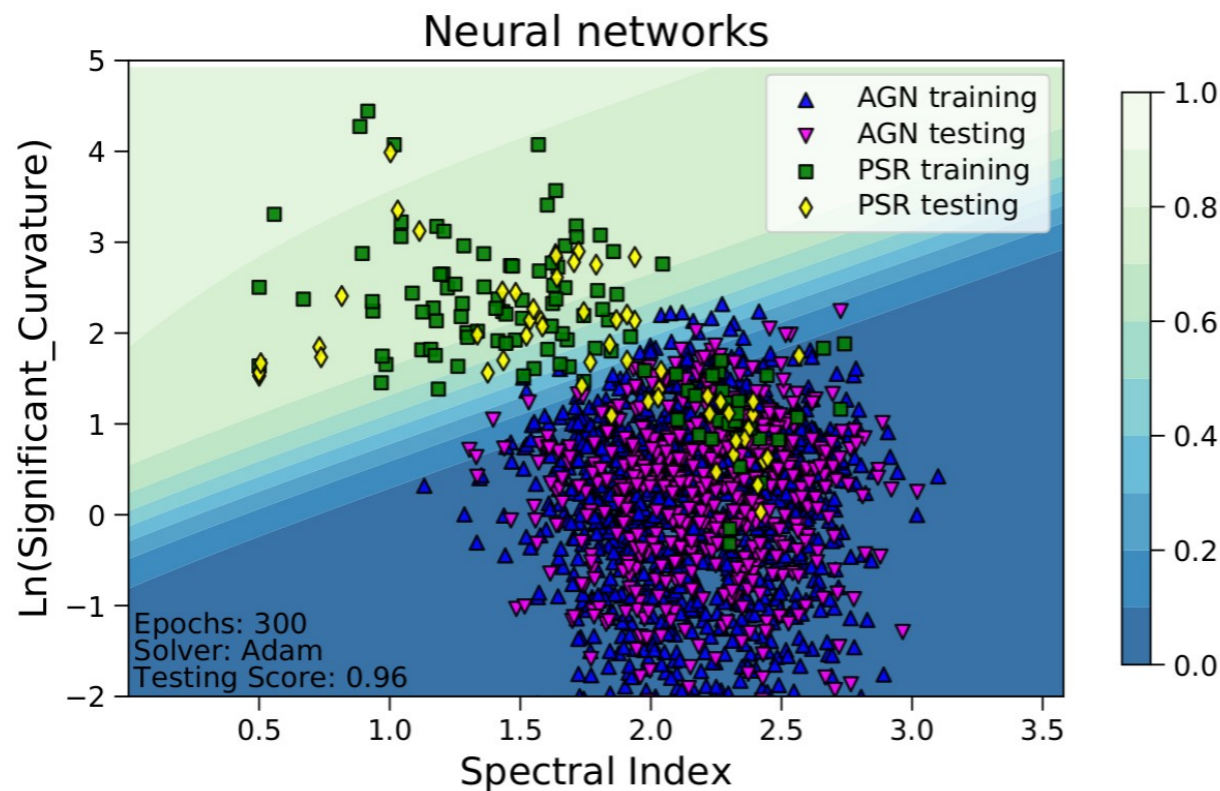
Isotropic background



Convolution: *Classic* versus *Graph* network



- Fermi LAT – gamma-ray astrophysics
 - **Aakash Bhat, Dmitry Malyshev**
 - Machine learning
 - Classification of unidentified point sources



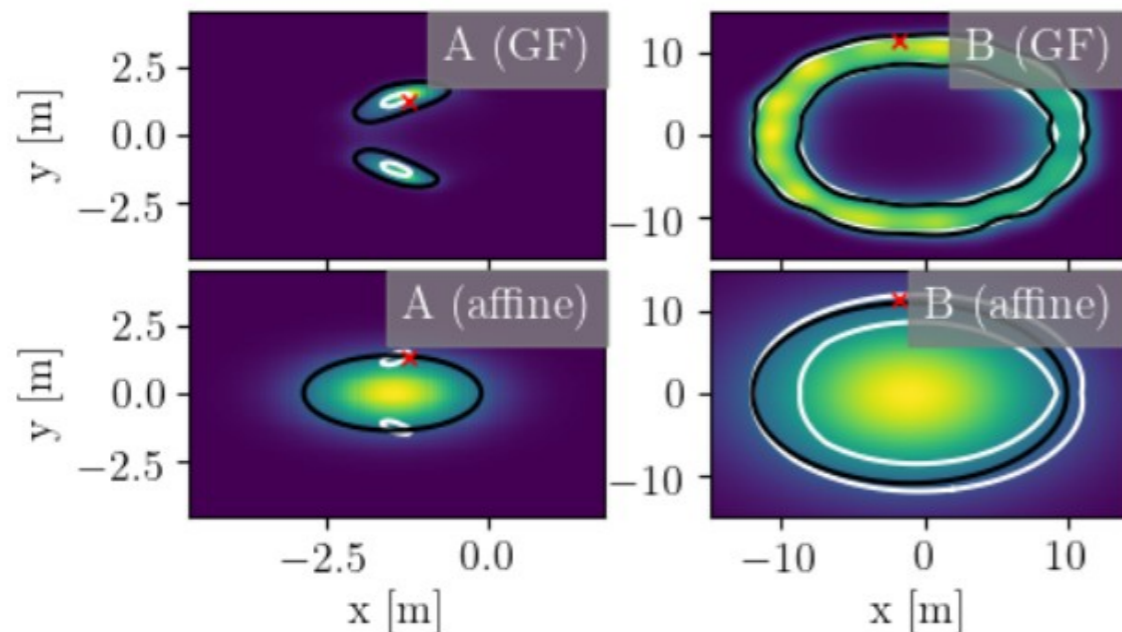
We use associated sources to train four ML algorithms (RF, BDT, LR, NN) to classify sources into active galactic nuclei (AGNs) and pulsars.

Normalizing flows → generalize supervised learning as approx. likelihood-free inference (2008.05825 – or see talk this meeting)

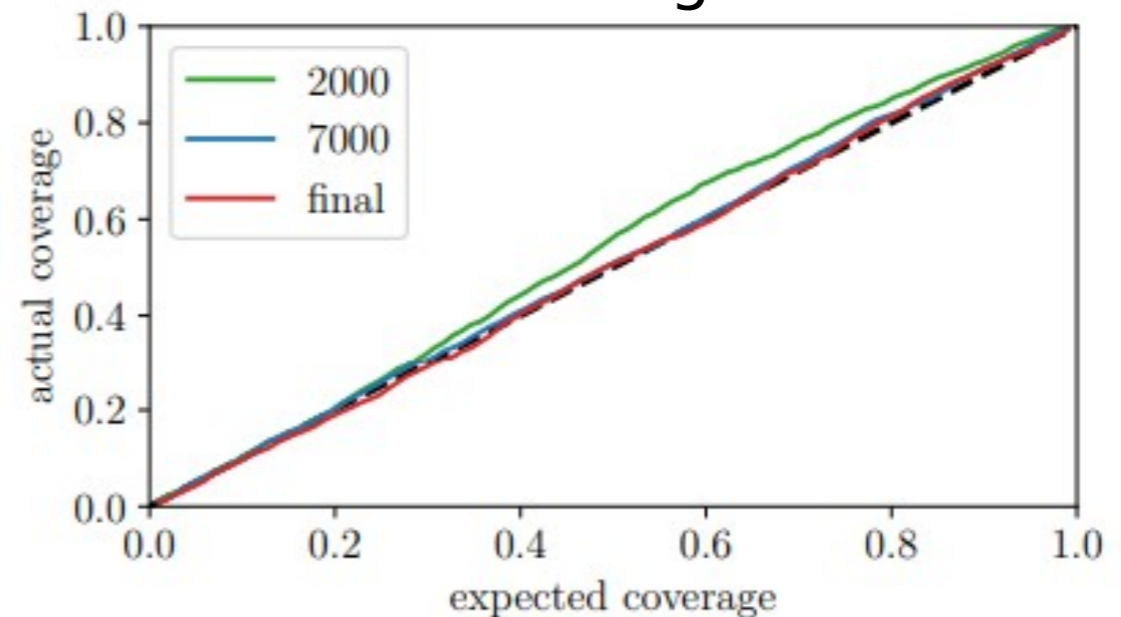
Thorsten Glüsenkamp

- 1) Systematics
 - 2) Goodness of fit
 - 3) Coverage
- for Neural network Posterior

True Posterior (White contour)
Approx Posterior (Black contour)



Coverage



Classification Comparison

- ErUM domains work with a diverse set of data representations
 - Different detector geometries, different types of experiments, theory calculations...
- But: several approaches should be flexible enough:
 - Fully connected
 - 1D convolutions
 - Graphs
 - Images
 - ...?
- Collecting datasets to enable cross-experiment / cross-disciplinary benchmarking of data representations and algorithms for our applications
- Focusing on classification as it allows straightforward benchmarking
- See: <https://github.com/erum-data-idt/ClassifierComparison>
(and contact me with more datasets/cross-dataset comparisons)

Datasets

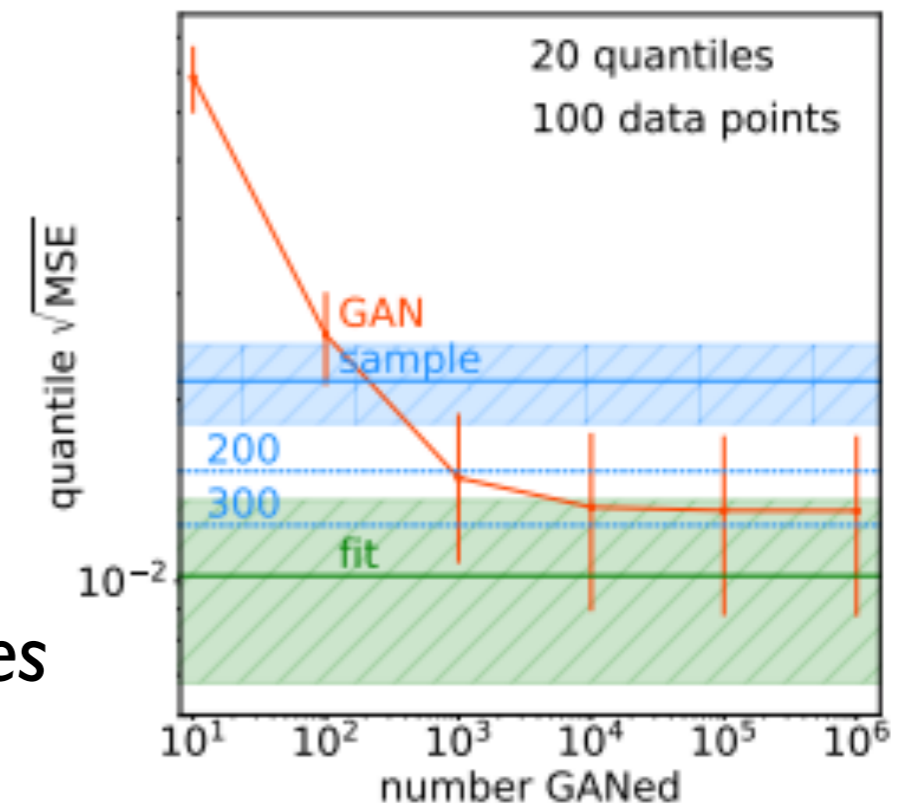
- Top Tagging at the LHC [link](#), Publication: 1902.09914
- Spinodal or not? [link](#), Publication: 1906.06562
- EOSL or EOSQ [link](#), Publication: 1910.11530
- Cosmic Airshower [link](#)
- LHC Olympics 2020 (Unsupervised anomaly detection) [link](#)

Fast generation (C3)

Hamburg

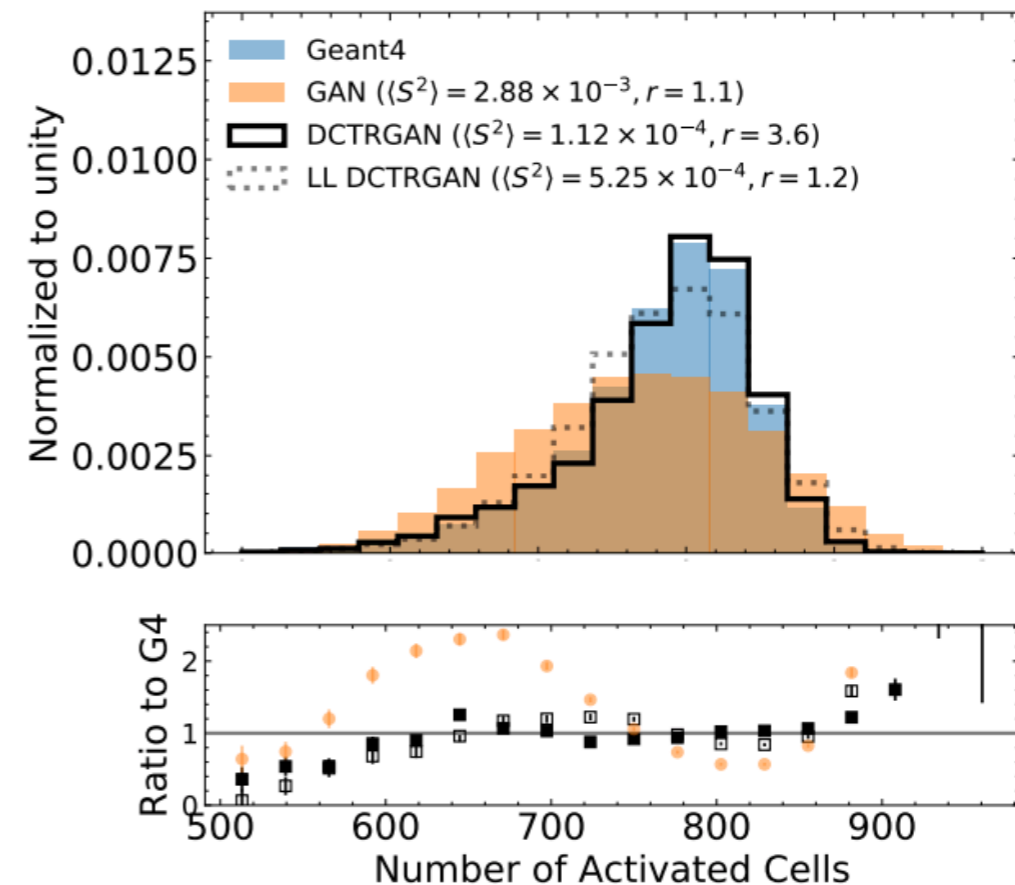
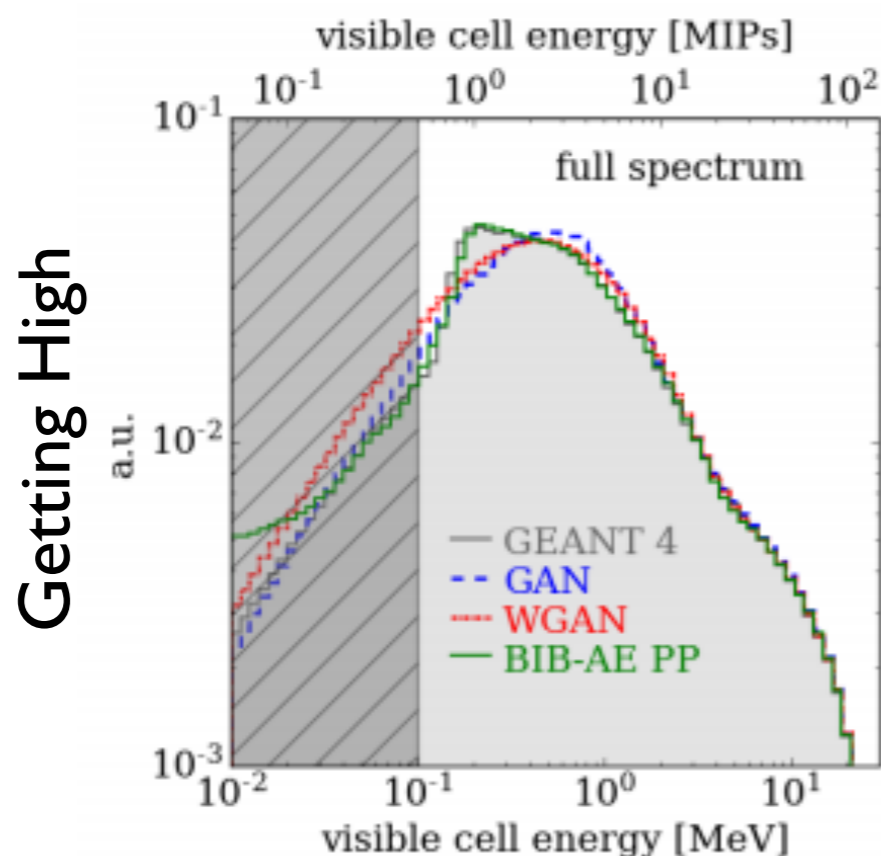
- Major motivation for generative models:
 - Speed-up event generation/simulation as they are computing bottle-necks
- How?
 - Generate N events using classical approach (e.g. Geant4)
 - Train generative model (e.g. GAN)
 - Sample M events from GAN
- Only helps if amplification ($M > N$) possible
- Sascha will present recent result tomorrow showing that this can be achieved for simple examples

arXiv:2008.06545 - *GANplifying Event Samples*



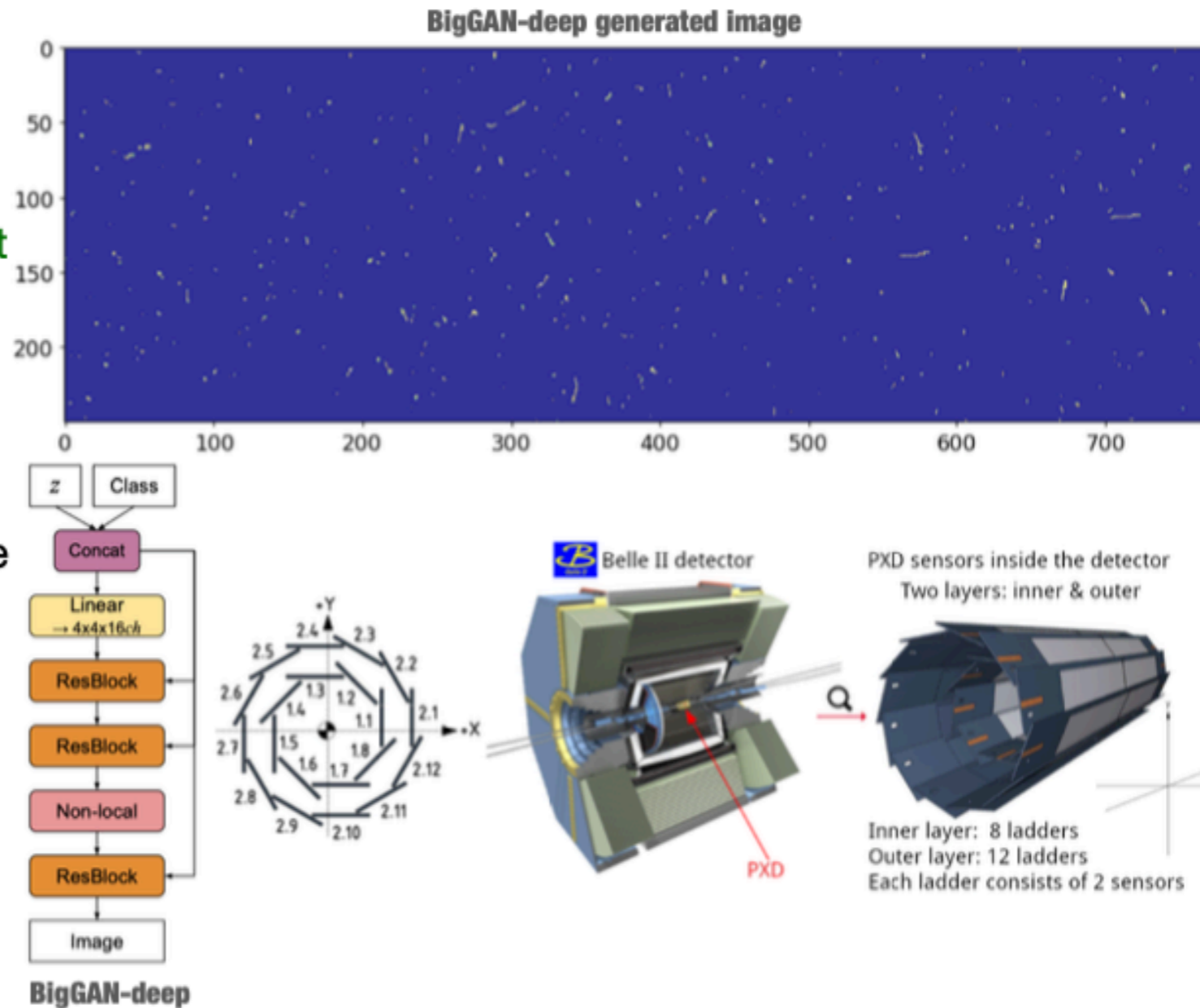
Hamburg

- Application: calorimeter showers for future colliders
- BIB-AE architecture achieves good fidelity
- update by Erik in last meeting, paper out now:
arXiv: 2005.05334 - *Getting High: High Fidelity Simulation of High Granularity Calorimeters with High Speed*
- Further improvement via event-wise re-weighting (refinement)
- Train classifier (full simulation vs GAN output) and use to derive weights
arXiv: 2009.03796 - *DCTRGAN: Improving the Precision of Generative Models with Reweighting*



Generating PXD-bg heat-maps using GAN

- The Pixel Vertex Detector (PXD) is the innermost semi-conductor sub-detector at Belle II.
- Detector simulation consists of both signal and background → PXD images have the highest storage consumption
- Using **GAN** instead of huge samples of Geant4 simulation or recorded Data.
- ✓ Initial Model: **WGAN-gp** New Model: **LOGAN**
- ✓ Using Class Conditional GAN for increasing image fidelity and generate sensor-based images: **BigGAN-deep**
- ✓ Validation of generated images:
 - Impact parameter resolution: **Frechet Distance**
 - Track reconstruction efficiency
- Outlook → Generate Images based on bkg type & capture correlation between layers



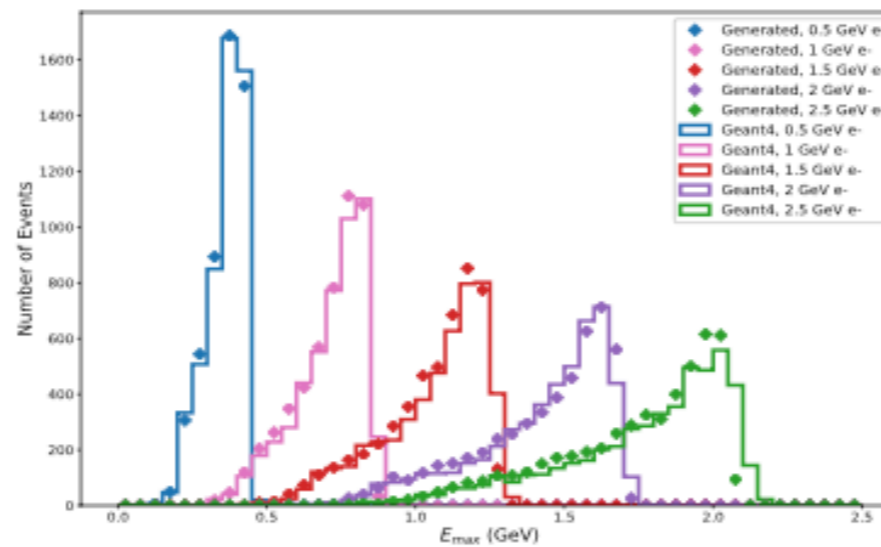
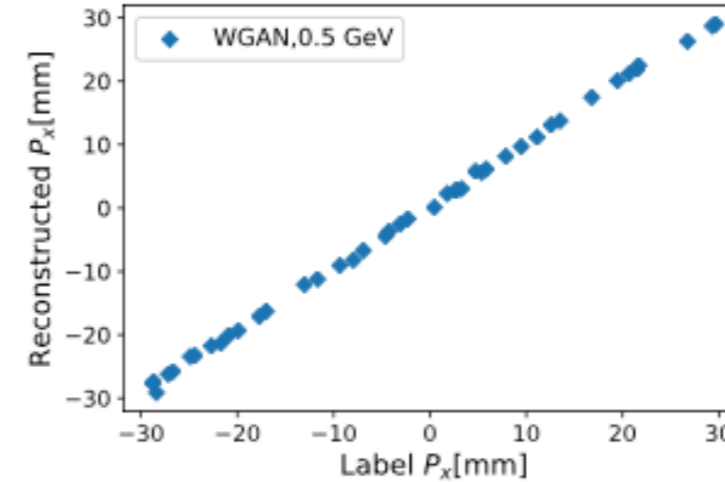
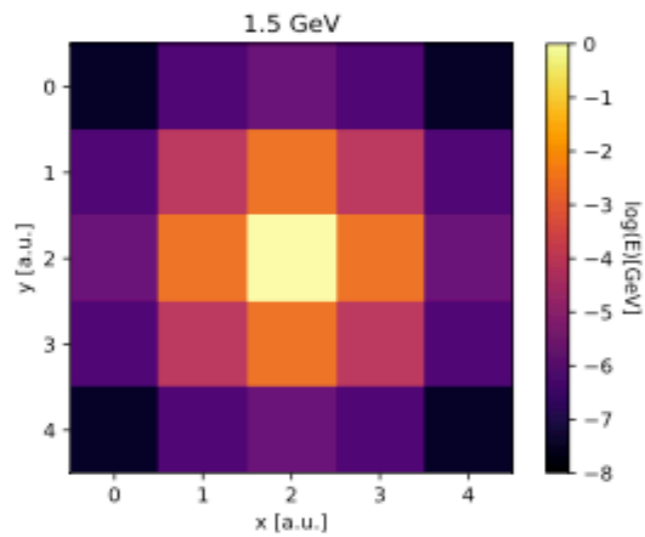
Fast Simulation of Particle Showers

Wasserstein GAN

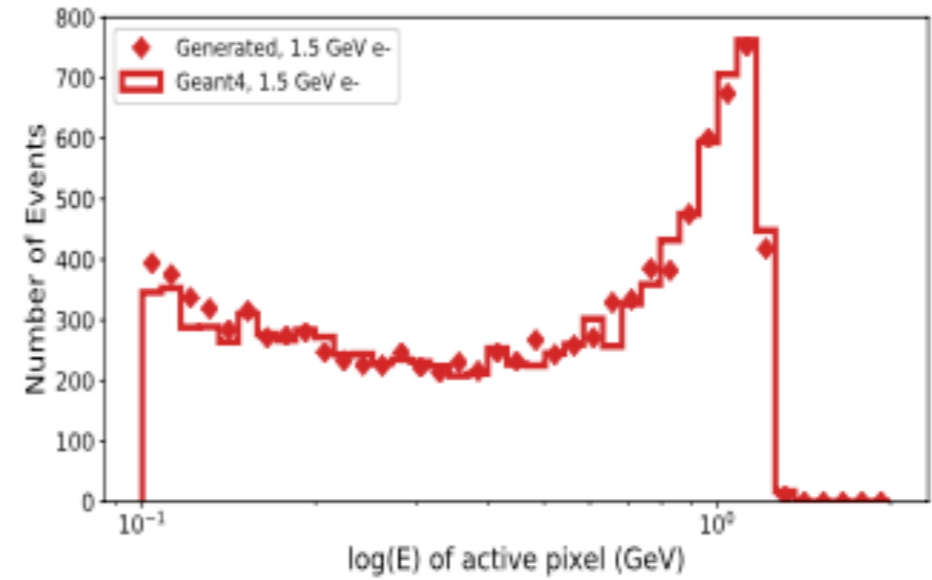
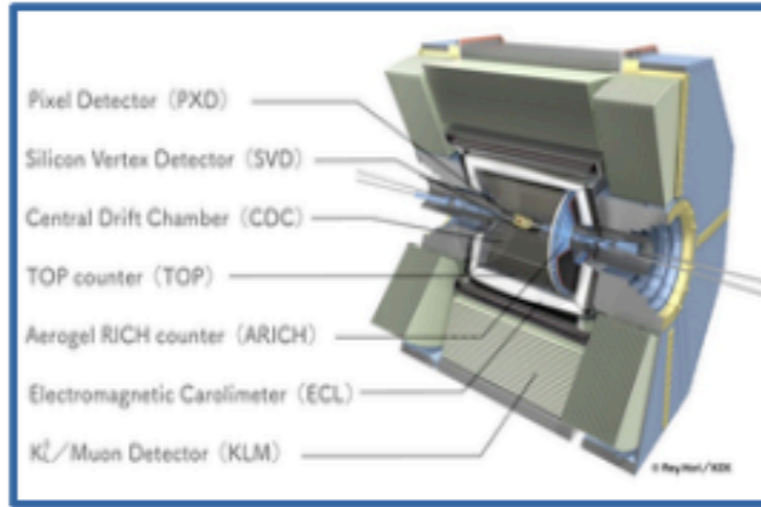


$$L_{\text{Critic}} = E[C(G(z))] - E[C(x)] + \lambda E[(\|\nabla_{\hat{x}} C(\hat{x})\|_2 - 1)^2]$$

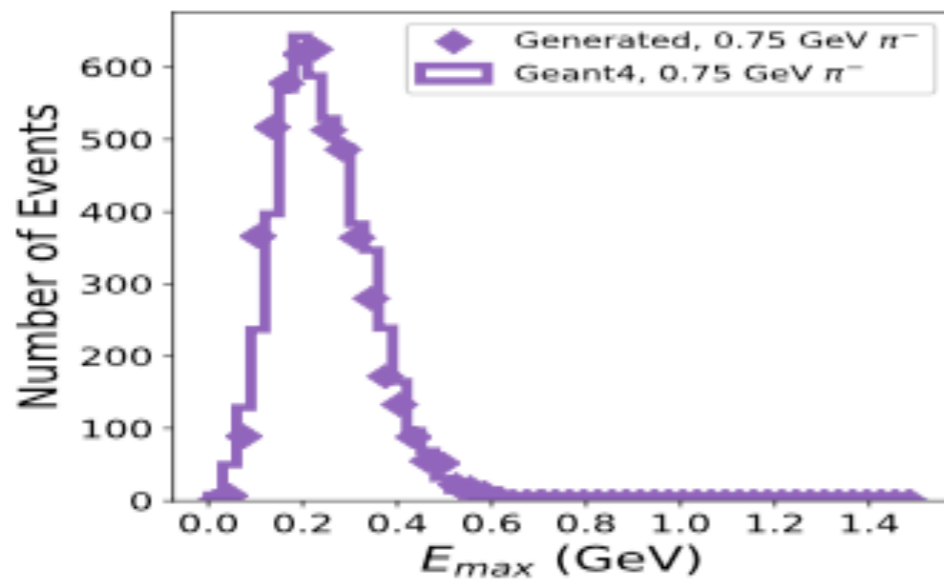
$$L_{\text{Generator}} = -E[C(G(z))]$$



Testing the model for electron showers in Belle II ECL crystals



Hadronic shower simulation

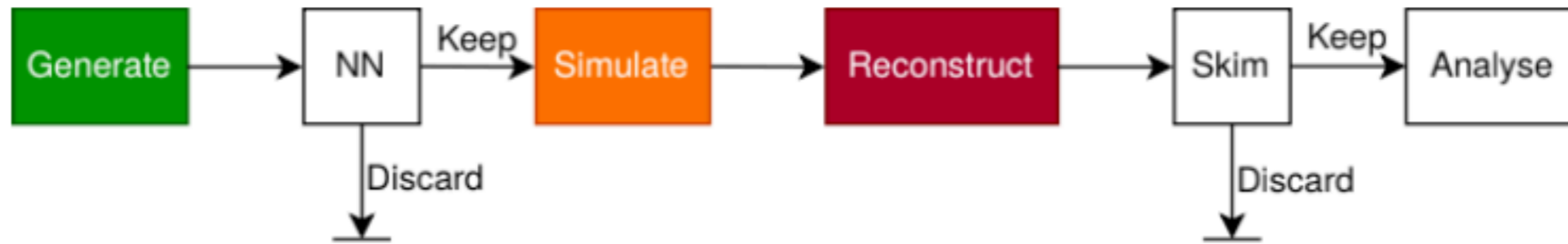


VAE-GAN

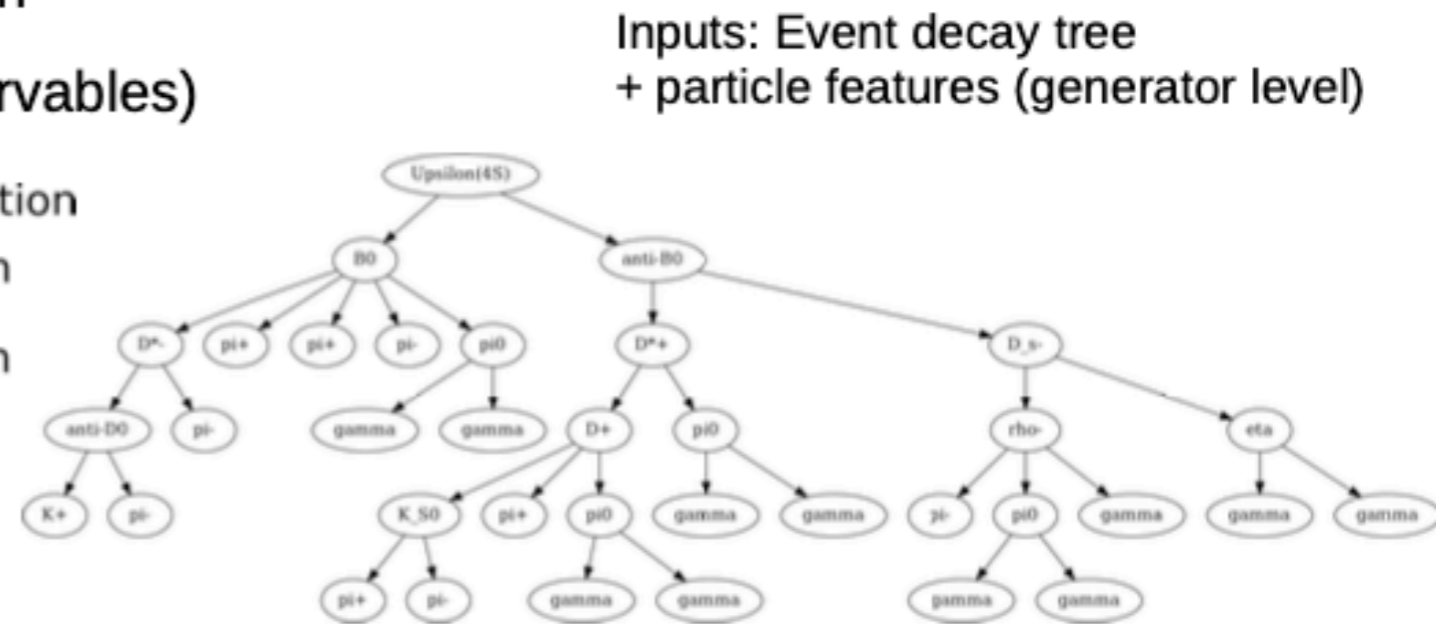
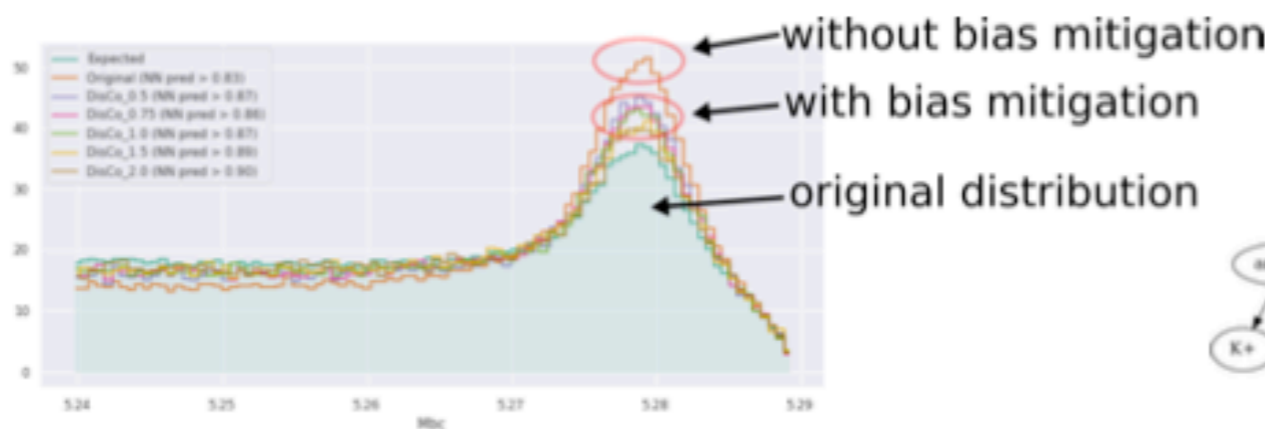


SmartBKG project at LMU

Try to filter Events already before expensive detector simulation and reconstruction



- Architectures based on graph convolutions have good performance
- Studied mitigation of bias due to false negatives
 - ➔ distance correlation term in loss function (decorrelate classifier output from observables)



- Preparing dataset for Classifier comparison

Closing

Summary

- Range of activities in ErUM-data C
- Algorithms suited to our data-structures
 - New results and convergence on similar approaches
 - New statistical tools for data analysis
- Machine learning close to hardware
 - Implementation of several key architectures done
 - The devil is in the details..
- Fast generative models
 - Many applications
 - First hints at possible amplification of statistics

Look forward to detailed updates tomorrow

Agenda for Tomorrow

Session D+C

D	09:00	15 - Status of ACTS Deployment at FASER	<i>Tobias Boeckh</i>
		16 - Status of Belle II ACTS Activities	<i>Ralf Farkas</i>
		17 - Impact parameter for heavy ion collisions with a PointNet architecture	<i>Manjunath Omana Kuttan</i>
C	10:00	24 - Identification of Cosmic Rays from Sources using Dynamic Graph Convolutional Neural Networks	<i>Niklas Langner</i>
		Short coffee break	
		18 - Using Generative Networks to amplify statistics	<i>Sascha Diefenbacher</i>
		19 - Flow-based networks and their benefits for high-energy physics	<i>Thorsten Glösenkamp</i>
	11:00	20 - Fast simulation of particle showers	<i>Jabbar Jubna</i>
		21 - Selective Background Monte Carlo simulation with deep learning	<i>Nikolai Hartmann</i>
		23 - PXD background image generation using GANs	<i>Hosein Hashemi</i>
	12:00		

Just updated!!!